DRAFT

12009

Olivine Vitrophyre

468.2 grams

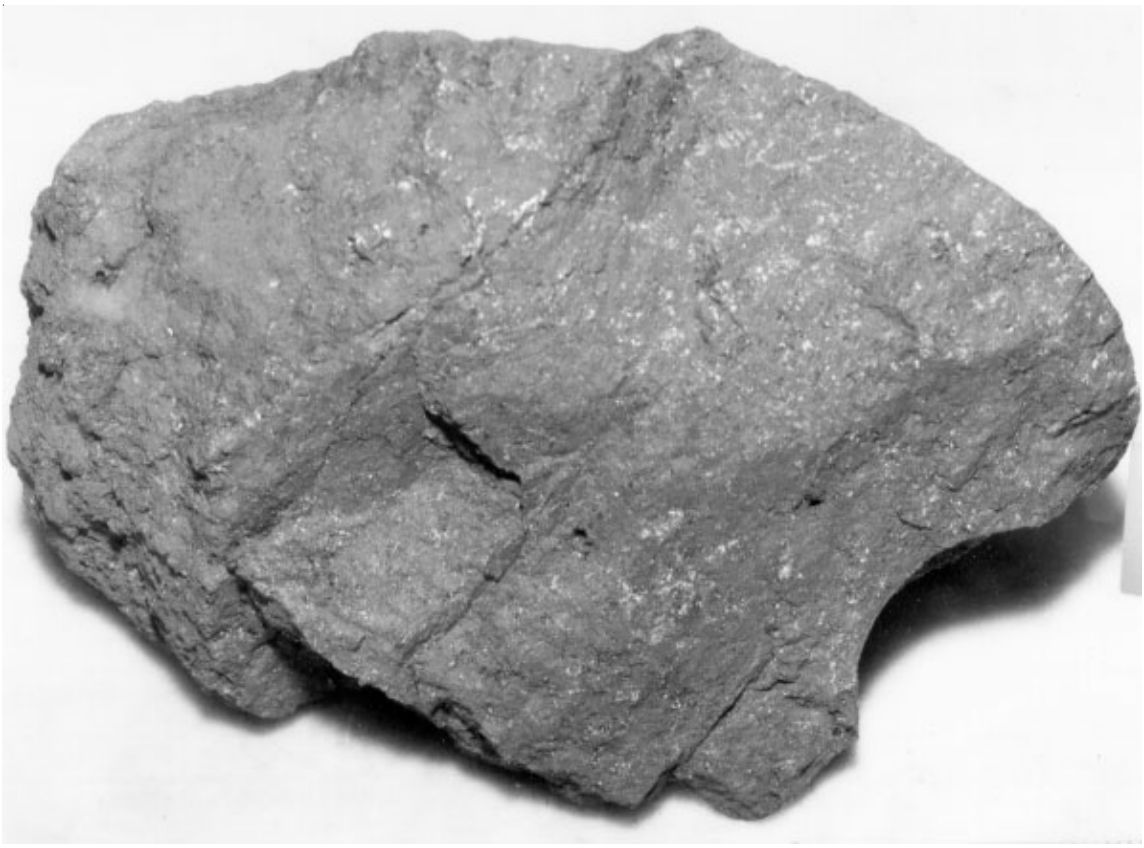


Figure 1: Photo of 12009,0 after dusting. Rock is 10 cm across. NASA photo # S70-47874.

Introduction

12009 is an olivine vitrophyre (quickly cooled basalt). It has portions of walls of large vesicles (figure 1). The collection site and field orientation of 12009 is not known.

Petrography

McGee et al. (1977) describe 12009 as "a porphyritic basalt vitrophyre which consists of skeletal phenocrysts of olivine (0.3 -1 mm) and pyroxene (0.2 – 0.8 mm) set in a matrix of microcrystalline devitrified glass and quench crystals of olivine, pyroxene and ilmenite". Drever et al. (1972) compared the texture of the skeletal olivine with that of selected terrestrial equivalents and Donaldson et al. (1975) studied the crystallization conditions that lead to this texture (figure 2, 3). Walker et al. (1976) discuss the crystal settling time needed to effect differentiation.

The Apollo 12 basalts can be related to one another by addition or subtraction of olivine (Fo_{74}) where the samples with the lowest mg* represent the liquid magma composition into which the liquidus olivine accumulated (Kushiro and Haramura 1971; Compston et al. 1971, James and Wright 1972, Green et al 1971 and Walker et al. 1976).

Mineralogy

Olivine: Beautiful skeletal phenocrysts of olivine are abundant in 12009 (Brett et al. 1971, Drever et al. 1972, Donaldson et al. 1975). Butler (1973) determined the minor element content of olivine. Walker et al. (1976) determined that the range of olivine composition was $Fo_{76\text{-}49}$.

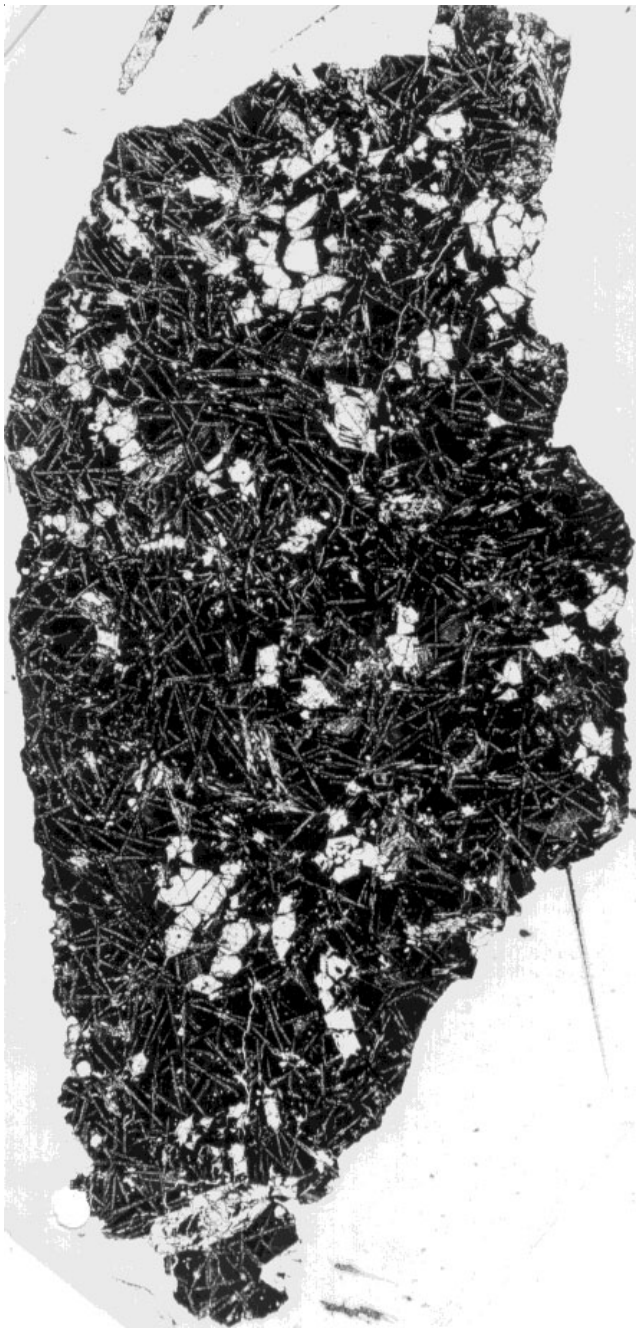


Figure 2: Photomicrograph of 12009,8 showing skeletal olivine phenocrysts in nearly opaque groundmass. NASA #S70-31568. Thin section is about 2 cm long.

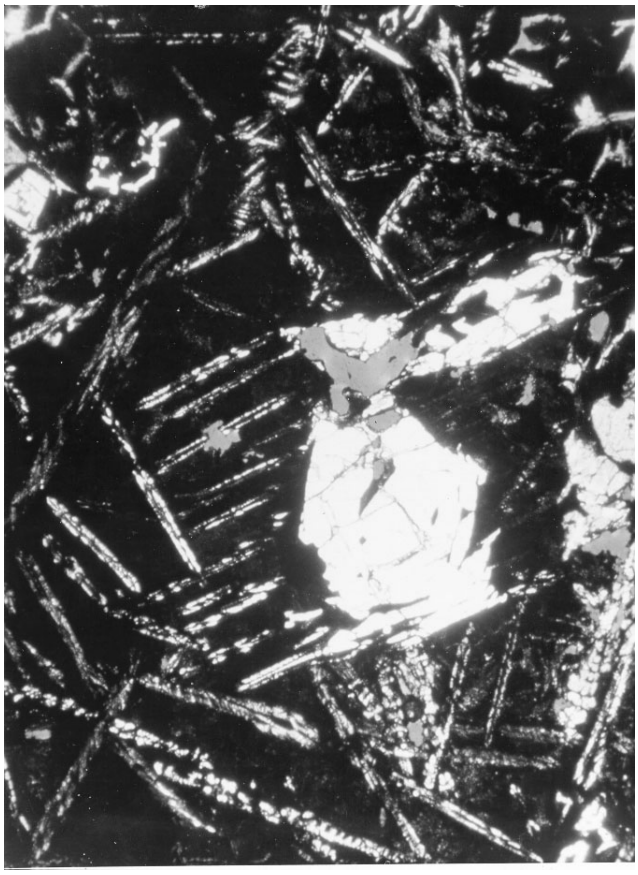


Figure 3: Photomicrograph of 12009,11 showing olivine phenocryst, olivine chains and feathery pyroxene in nearly opaque groundmass. NASA # S70-49828. Field of view is 2.2 mm.

Pyroxene: Pyroxene phenocrysts occur as bundles of elongate fibrous crystals that are often in optical continuity. Groundmass pyroxene is feathery. McGee et al. (1977) determined pyroxene composition (figure 4).

Metal: Brett et al. (1971) determined the Ni content of minute metallic iron grains in 12009 (figure 5).

Mineralogical Mode for 12009

	Neal et al. 1994	Brett et al. 1971	Papike et al. 1976
Olivine	48.8	23.6	21.7
Pyroxene		3.8	9.9
Plagioclase	0.1		
Ilmenite	1.2	0.3	0.6
Chromite +Usp	2.2		
mesostasis	47.2	72.3	68

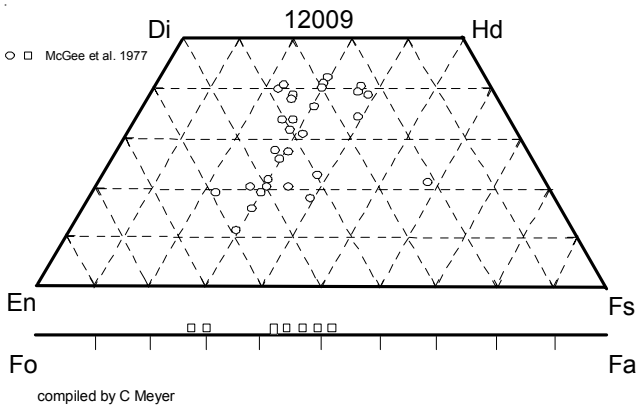


Figure 4: Pyroxene and olivine composition of 12009 (adapted from McGee et al. 1977).

Chemistry

The chemical composition of 12009 seems to be the same as for 12015 (Table 1, figure 6 and 7). It has often been used in modeling the origin and differentiation of Apollo 12 basalts.

Radiogenic age dating

Stettler et al. (1973) determined an age for 12009 by total argon 39/40 (see table). The high temperature release represented a lower age (see figure in 12051). Snyder et al. (1997) reported the isotopic composition of Sr and Nd.

Cosmogenic isotopes and exposure ages

Stettler et al. (1973) determined an ^{38}Ar exposure age of 140-160 m.y. Marti and Lugmair (1971) determined an ^{81}Kr –^{83}Kr exposure age of 136 ± 24 m.y.

Other Studies

Bogard et al. (1971) reported the content and isotopic composition of rare gases in 12009. Green et al. (1971) studied the high pressure phase equilibria for 12009 (figure 8). Donaldson et al. (1975) studied the cooling history

Processing

12009 was broken to create subsamples (see figure).

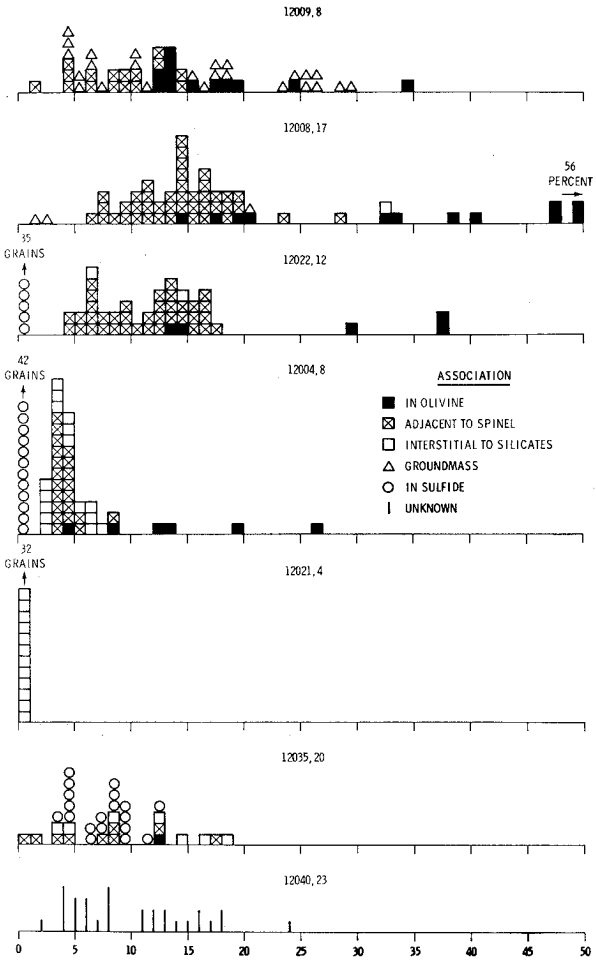


Figure 5: Histogram of Ni conentrations of metal grains in 7 lunar samples (lifted from Brett et al. 1971).

List of Photo #s for 12009

S69-62296-62307	B&W mug
S69-62739-62743	B&W
S69-64090-091	color
S69-64115-116	
S70-31568	TS ,8
S70-47870-874	processing
S70-49151-154	TS
S70-49244-251	TS
S70-49827-828	TS
S75-32779-780	color
S79-27094-093	TS

Summary of Age Data for 12009

	Ar/Ar	Rb/Sr
Stettler et al. 1973	3.29 ± 0.07 b.y.	
	3.17 ± 0.07	

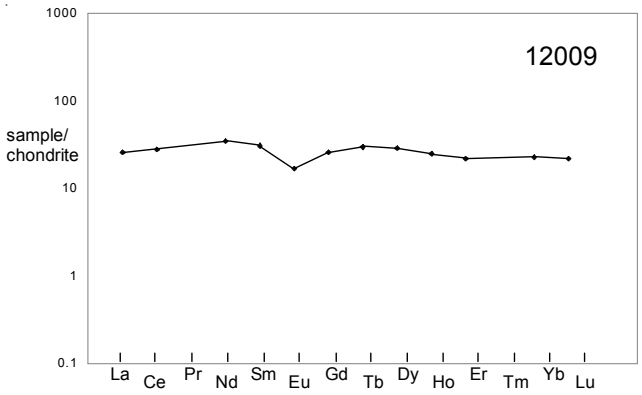


Figure 6: Normalized rare-earth-element diagram for 12009 (data from Haskin et al. 1971).

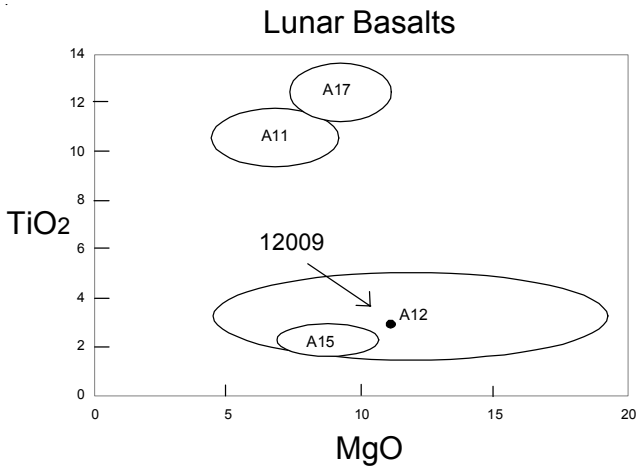


Figure 7: Composition of 12009 compared with that of lunar basalts.

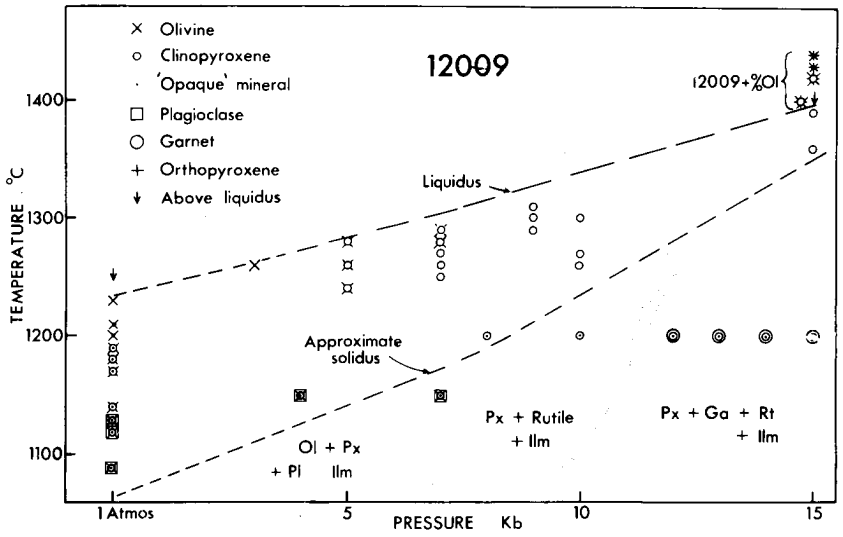
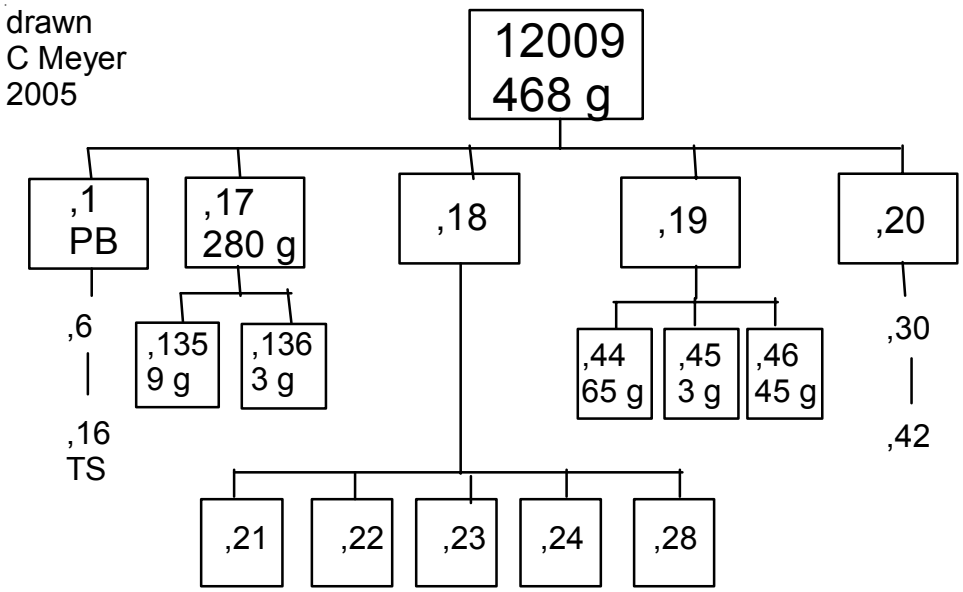


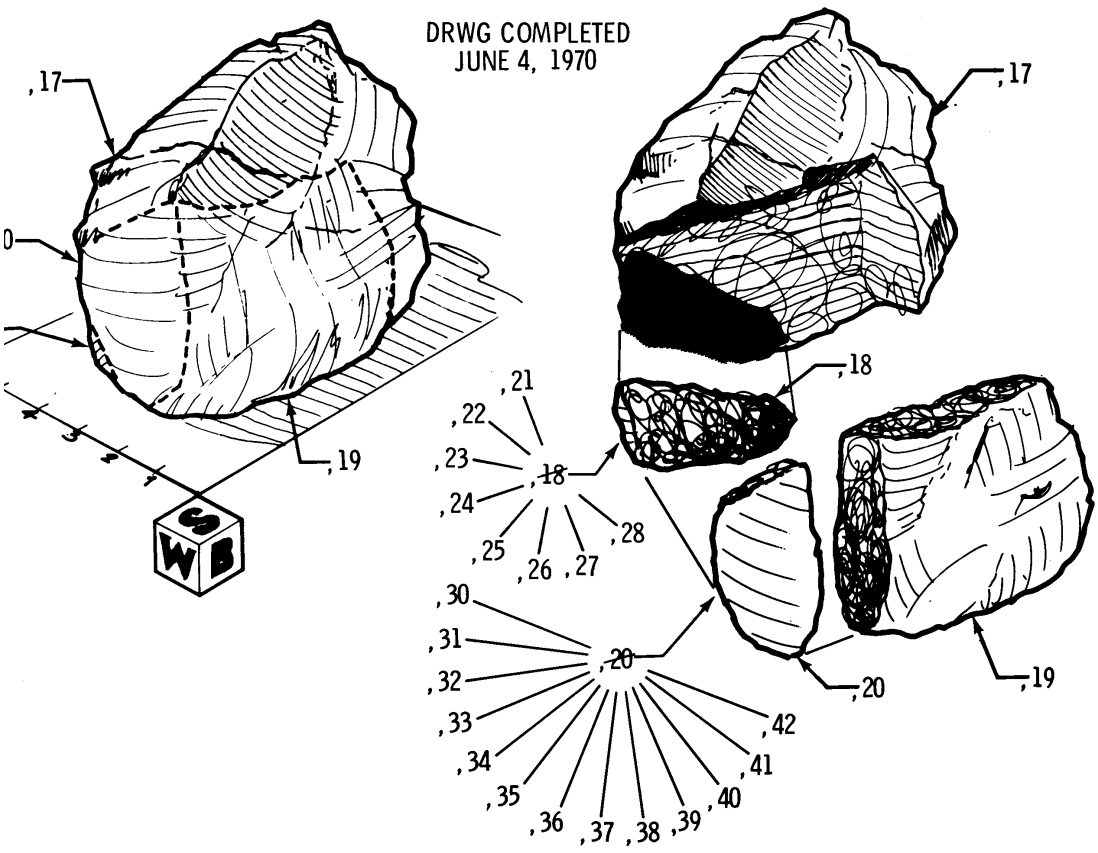
Figure 8: Experimental phase diagram for liquids with 12009 composition as function of temperature and depth of origin (Green et al. 1971).

Table 1. Chemical composition of 12009.

reference	LSPET70	Murthy71		Cuttitta71		Haskin71		Compston71		Tats71		Baedecker71		Snyder97	
weight															
SiO2 %	41							45.03	(d)					45	
TiO2	3.3							2.9	(d)					2.9	
Al2O3	11							8.59	(d)					8.59	
FeO	20							21.03	(d)					21	
MnO	0.19			0.27	(b)			0.28	(d)					0.28	
MgO	12.5							11.55	(d)					11.6	
CaO	10							9.42	(d)					9.42	
Na2O	0.51							0.23	(d)					0.23	
K2O	0.063	0.047	(a)					0.064	(d)					0.06	
P2O5								0.07	(d)					0.07	
S %								0.06	(d)						
sum															
Sc ppm	42			44	(b)									46	(f)
V	77			166	(b)			153	(d)						
Cr	5200			4390	(b)			3790	(d)					3960	(f)
Co	46			59	(b)			49	(d)					50.1	(f)
Ni	67			62	(b)			52	(d)					55	(f)
Cu				14	(b)									10.4	(f)
Zn				4	(b)							1.8	(e)	9.7	(f)
Ga				5	(b)			2	(d)			3.2	(e)	3.11	(f)
Ge ppb															
As															
Se															
Rb	0.57	1.09	(a)	1.4	(b)			1.03						0.987	(f)
Sr	110	89.8	(a)	75	(b)			95.6						86.4	(f)
Y	48			42	(b)			34						36.5	(f)
Zr	150			114	(b)			107						106.4	(f)
Nb								6						6.5	(f)
Mo															
Ru															
Rh															
Pd ppb															
Ag ppb														82	(f)
Cd ppb												2.2	(e)		
In ppb												1.6	(e)		
Sn ppb															
Sb ppb															
Te ppb															
Cs ppm														0.068	(f)
Ba	65	76	(a)	71	(b)			60						55.5	(f)
La						6.1	(c)	4						5.62	(f)
Ce						16.8	(c)	10						16.1	(f)
Pr														2.45	(f)
Nd						16	(c)							12.7	(f)
Sm						4.53	(c)							3.91	(f)
Eu						0.94	(c)							0.89	(f)
Gd						5.2	(c)							4.15	(f)
Tb						1.11	(c)							0.9	(f)
Dy						7.13	(c)							5.7	(f)
Ho						1.4	(c)							1.14	(f)
Er						3.6	(c)							3.39	(f)
Tm														0.48	(f)
Yb				5.3	(b)	3.74	(c)							3.05	(f)
Lu						0.55	(c)							0.45	(f)
Hf															
Ta														0.296	(f)
W ppb															
Re ppb															
Os ppb															
Ir ppb												0.08	(e)		
Pt ppb															
Au ppb															
Th ppm										0.881	(a)			0.85	(f)
U ppm										0.243	(a)			0.23	(f)

technique: (a) IDMS, (b) mixed, (c) INAA, (d) XRF, (e) RNAA, (f) ICP-MS


drawn
C Meyer
2005
12009
468 g
,1
PB
,17
280 g
,18
,19
,20
,6
,16
TS
,135
9 g
,136
3 g
,44
65 g
,45
3 g
,46
45 g
,30
,42
,21
,22
,23
,24
,28


DRWG COMPLETED
JUNE 4, 1970
,17
,19
,18
,20
,21
,22
,23
,24
,25
,26
,27
,28
,30
,31
,32
,33
,34
,35
,36
,37
,38
,39
,40
,41
,42
S
W
B